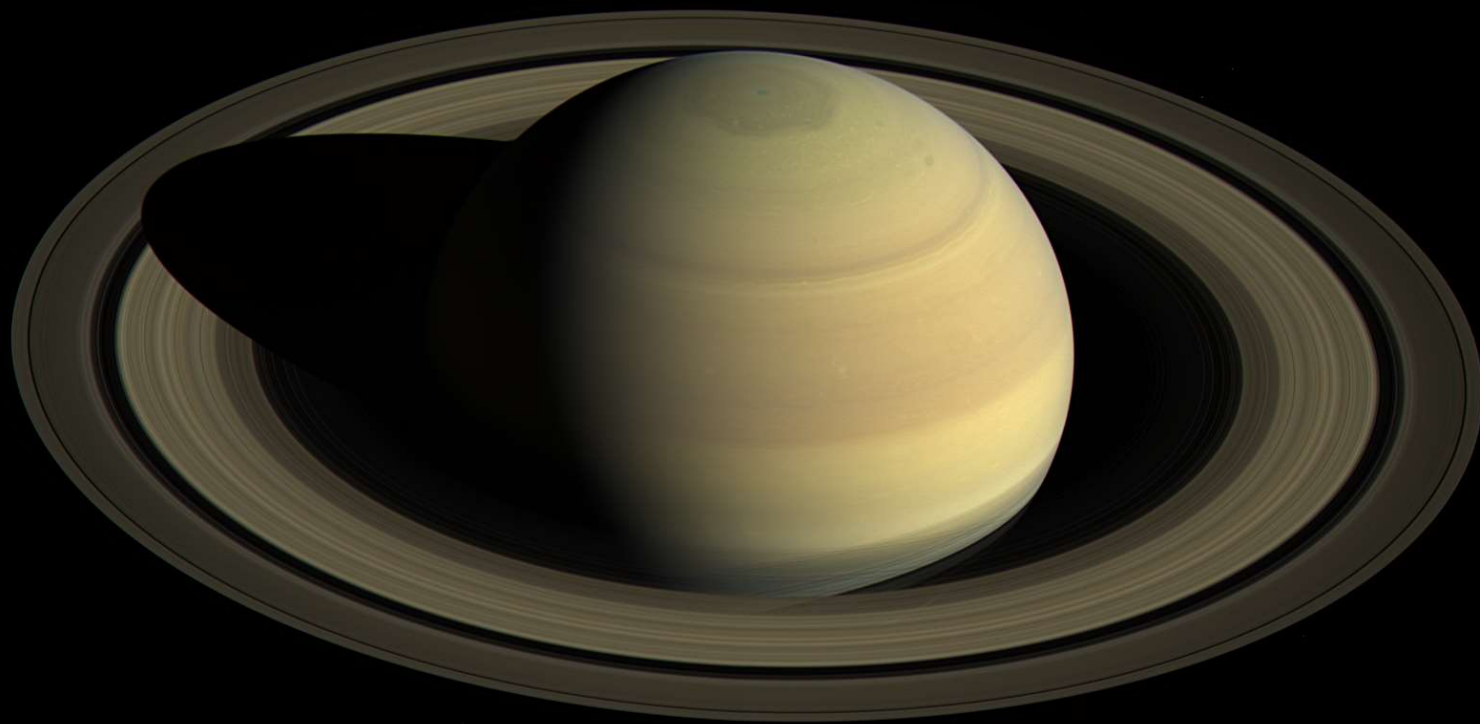


# **AST 2002**

## **Introduction to Astronomy**



# iClicker2

**Available at the bookstore.  
At the beginning of every class  
will sync with iClicker base  
(code will be BC).**



**Must register on Webcourses using knights email  
address within first week of class**

**Participation counts for 5%**

**Correct answers for 5%**

**We will use the 1<sup>st</sup> week to practice taking polls**

**Material may be based on reading material to be  
covered that lecture**

# iClicker2 Instructions

1. Press and hold the Power/Change Frequency button until the two-letter frequency on the LCD flashes.
2. Use the A-E buttons to enter the new two-letter frequency code. Enter 'BC' as the code. A checkmark appears on the LCD indicating the frequency change was successful. It should welcome you by showing "AST2002".
3. When you supply an answer, you should be able to see a ✓

## Question #1

What is the correct order, from smallest to largest?

- A: An Asteroid, The Milky Way Galaxy, The Solar System, The Universe
- B: The Universe, The Milky Way Galaxy, An Asteroid, The Solar System
- C: An Asteroid, The Solar System, The Milky Way Galaxy, The Universe
- D: An Asteroid, The Solar System, The Universe, The Milky Way Galaxy

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3. When you supply an answer, you should be able to see a ✓

## Question #2

What is the correct order, from smallest to largest?

- A: A Jovian Planet, A Terrestrial Planet, One Astronomical Unit (AU), a Light-year
- B: A Terrestrial Planet, A Jovian Planet, One Astronomical Unit (AU), a Light-year
- C: A Terrestrial Planet, a Light-year, One Astronomical Unit (AU), A Jovian Planet
- D: A Jovian Planet, A Terrestrial Planet, a Light-year, One Astronomical Unit (AU)

# iClicker2 Instructions

1. Press and hold the Power/Change Frequency button until the two-letter frequency on the LCD flashes.
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## Question #2

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- A: A Jovian Planet, A Terrestrial Planet, One Astronomical Unit (AU), a Light-year
- B: A Terrestrial Planet, A Jovian Planet, One Astronomical Unit (AU), a Light-year**
- C: A Terrestrial Planet, a Light-year, One Astronomical Unit (AU), A Jovian Planet
- D: A Jovian Planet, A Terrestrial Planet, a Light-year, One Astronomical Unit (AU)

# We Would Like to Understand This Better... ... by the end of Today's lecture, you will

The patterns we observe in the night sky are affected by:

- The rotation of the Earth
- The revolution of the Earth around the Sun
- The axial tilt of the Earth
- The movement of the Moon around the Earth
- The movement of the planets around the Sun



2017 Jan 31 Tue, 1 hour after sunset  
6:19 PM CST = Feb 1, 0:19 UT  
latitude 40°N, longitude 90°W  
sidereal time 46° = 3.08h

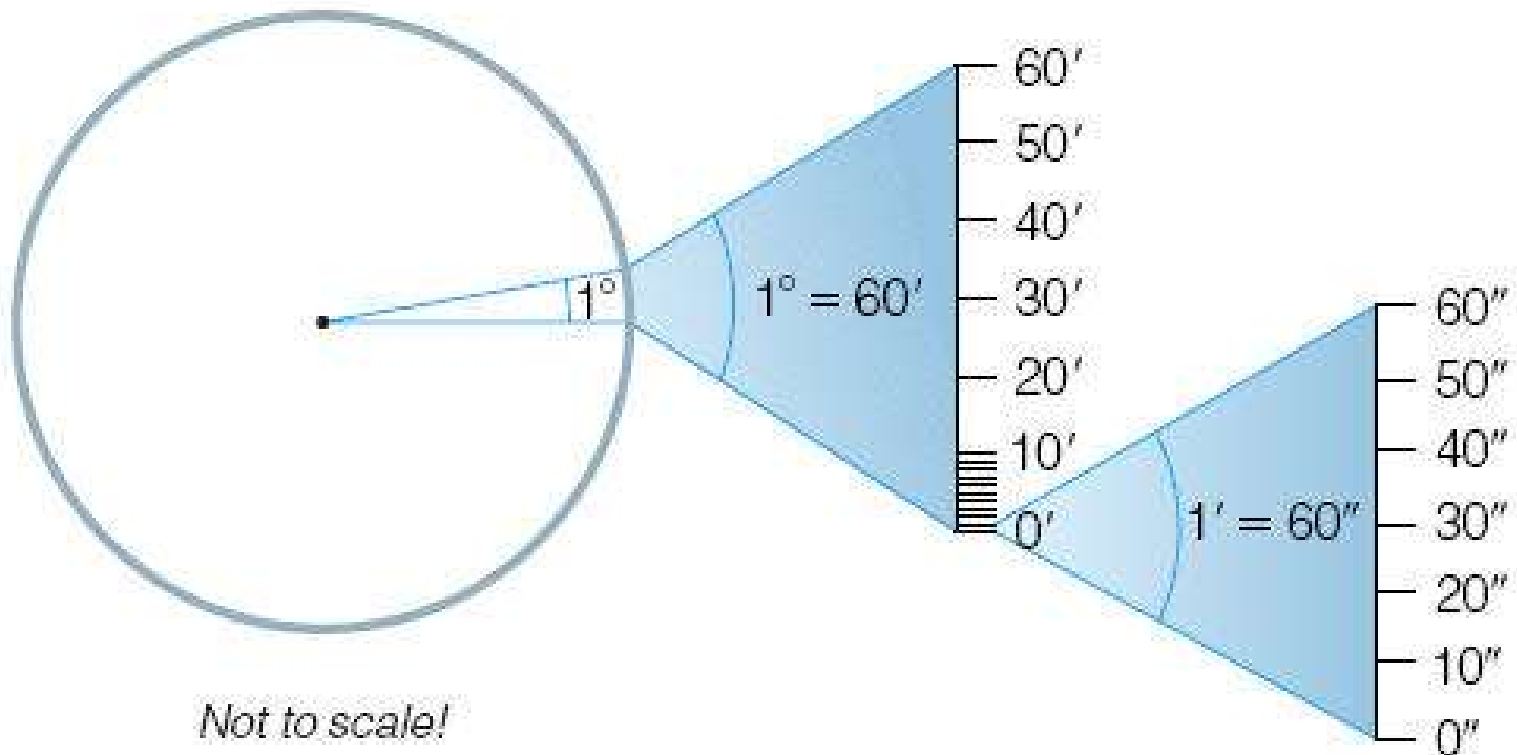
# We measure the sky in *angles*

Full circle =  $360^\circ$  (degrees of arc)

$1^\circ = 60'$  (arcmin)

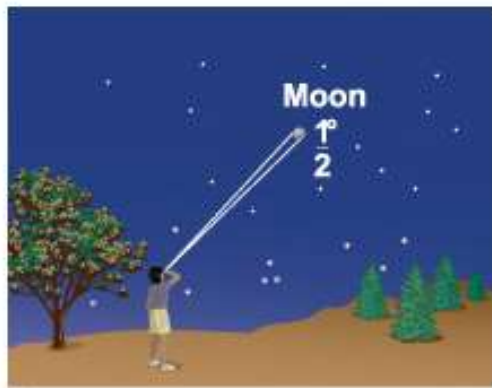
1 arcmin =  $60''$  (arcsec)

(May also see  $360^\circ \sim 24$  hours)

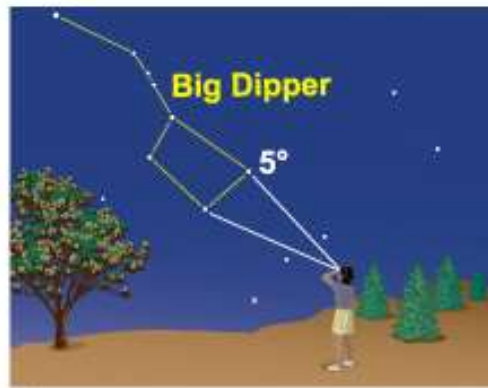




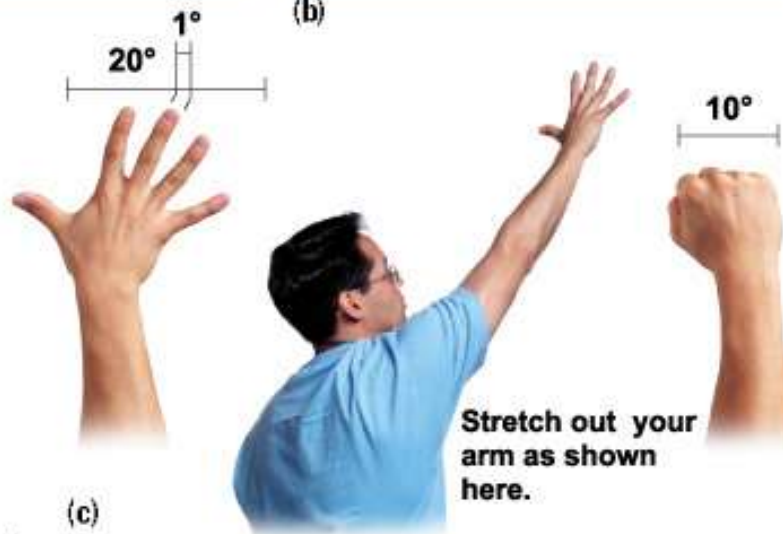
# How to Approximately Measure Angles in the Sky



(a)

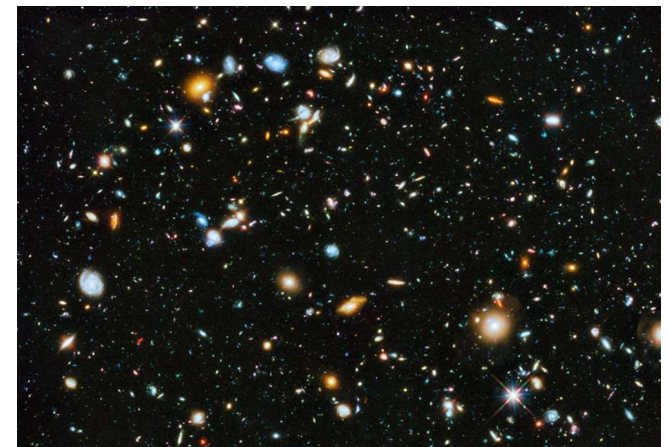


(b)



(c)

Hubble Ultra-Deep Field



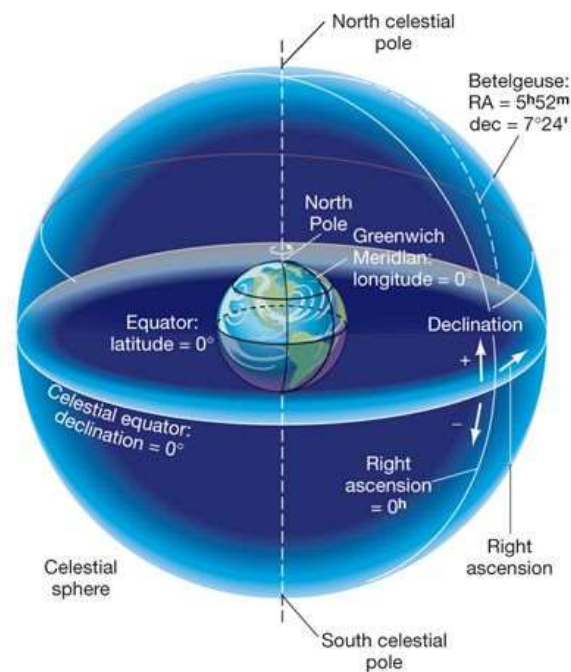
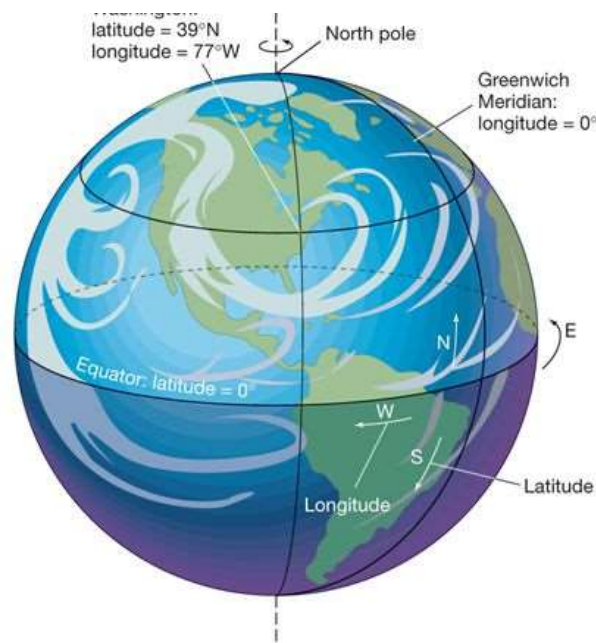
Width or Height = 2.4' arcminutes

roughly one thirteen-millionth of the total area of the sky

# Describing Positions on the Celestial Sphere

**Declination:** Degrees north or south of celestial equator (*cf. Latitude*)

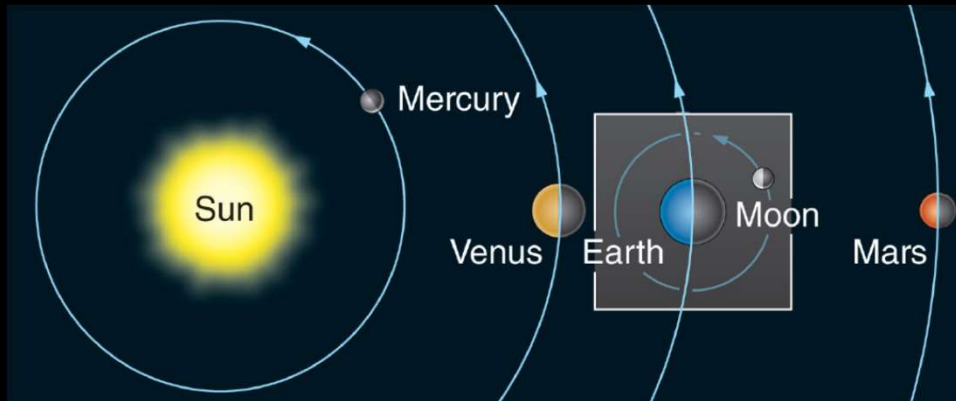
**Right Ascension:** Measured eastward position from the  $0^\circ$  (given in degrees, or hours and minutes).  $0^\circ$  is defined as the position of the Sun at *Vernal equinox* (*cf. longitude*)



# Solar Day vs. Sidereal Day (take II)

<http://www.jgiesen.de/astro/astroJS/siderealClock/>  
<http://aa.usno.navy.mil/data/docs/siderealtime.php>

Remember: It takes Earth 365 days to rotate 360°



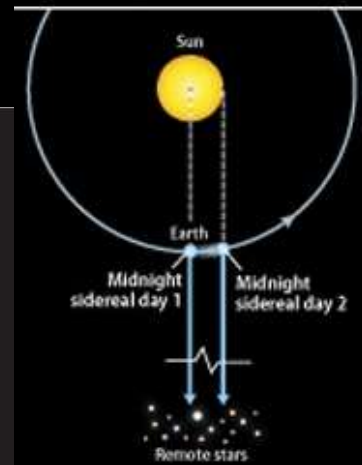
## Sidereal day

Earth rotates 360° in 23 hours, 56 minutes, 4 seconds

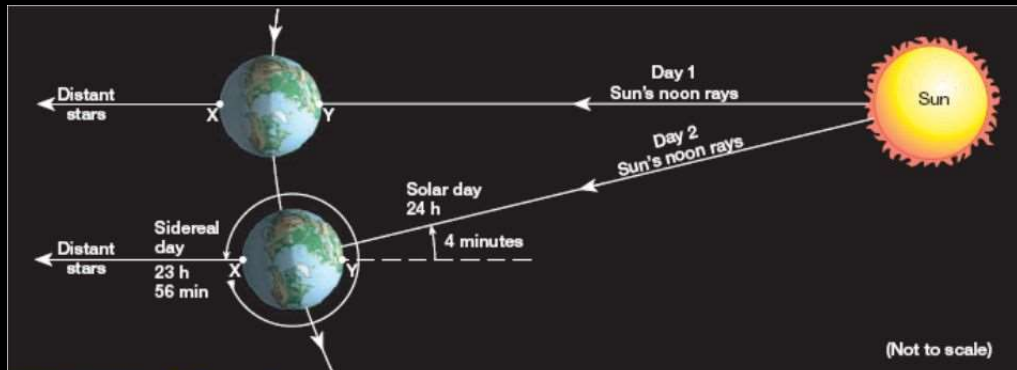


## Solar day

Earth rotates 360° + 1° in 24 hours



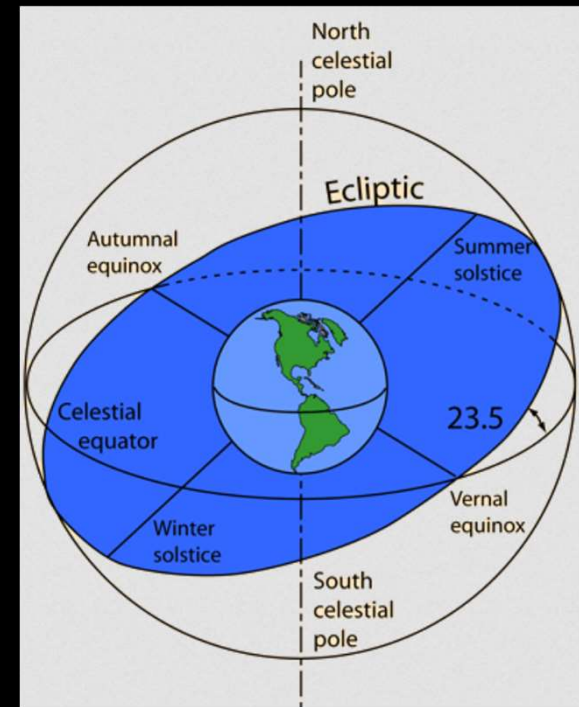
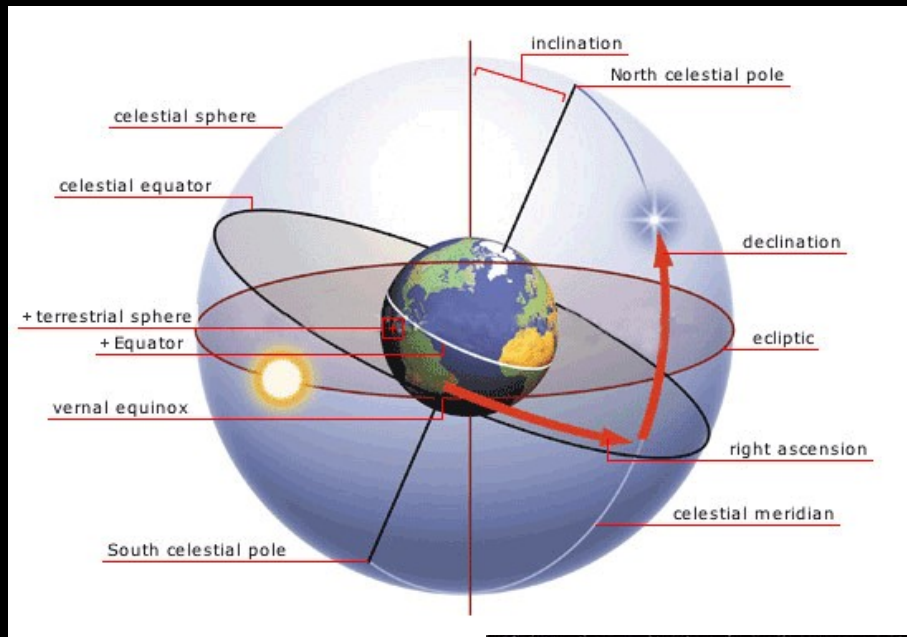
Not to scale  
Rotation angles and sizes exaggerated to show concept



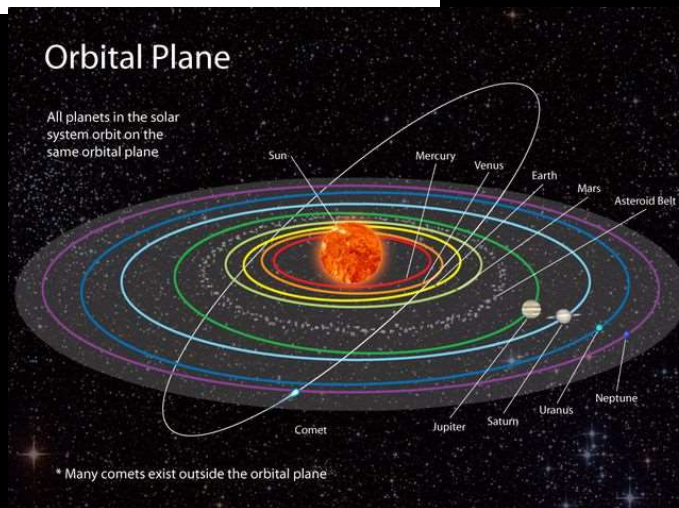


# Earth is Tilted 23.5° to the Ecliptic Plane...

The Sun, the Moon, the planets and the constellations of the zodiac all lie in the Ecliptic plane (orbital plane of solar system)



**Viewed from Ecliptic Plane:**  
Earth axis is tilted if viewed in ecliptic plane, celestial equator tilted 23.5°



**Viewed from Celestial Plane:**  
If image is aligned with North celestial pole and equator, and Earth is not tilted

# Annual Motion

**Ecliptic:** the apparent path of the Sun, moon, and planets through the sky

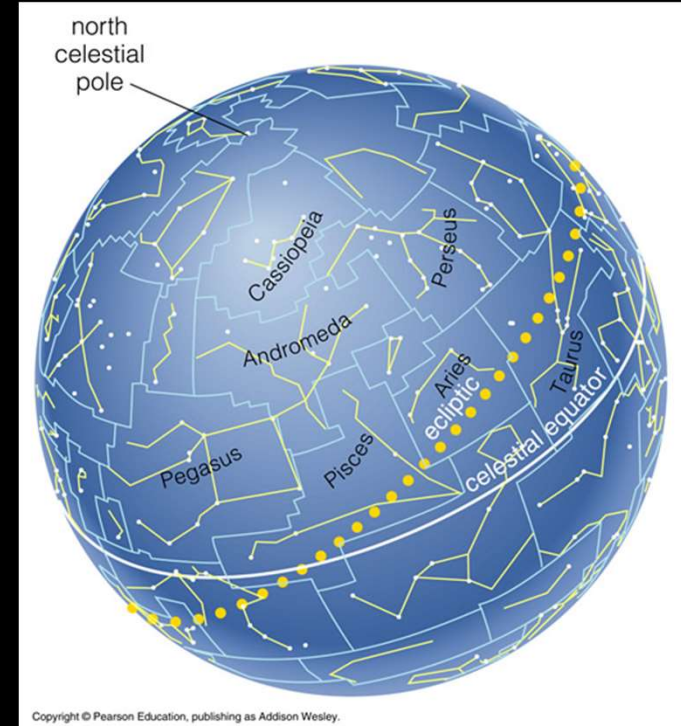
**Equinox:** where the ecliptic intersects the celestial equator (days and nights ~ equal)

## Solstice

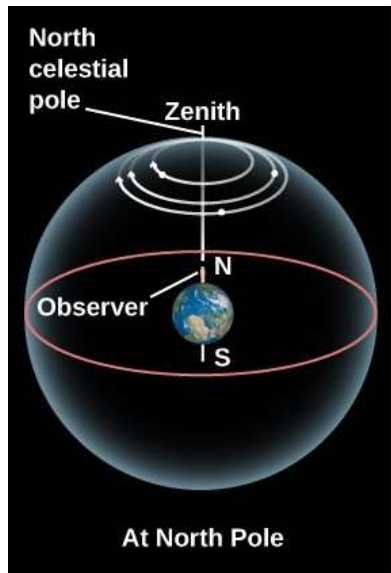
where the ecliptic is farthest from the celestial equator (Sun, stars of particular constellations reach max/min point)

## Zodiac

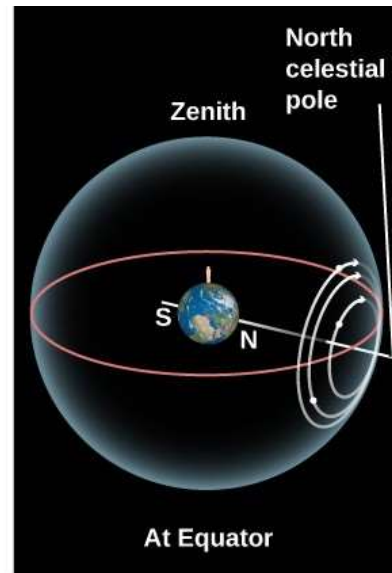
the constellations which lie along the ecliptic



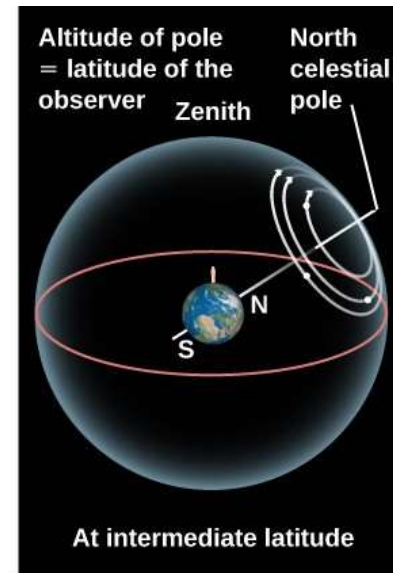
# How does the night sky change by location on the Earth?



(a)



(b)

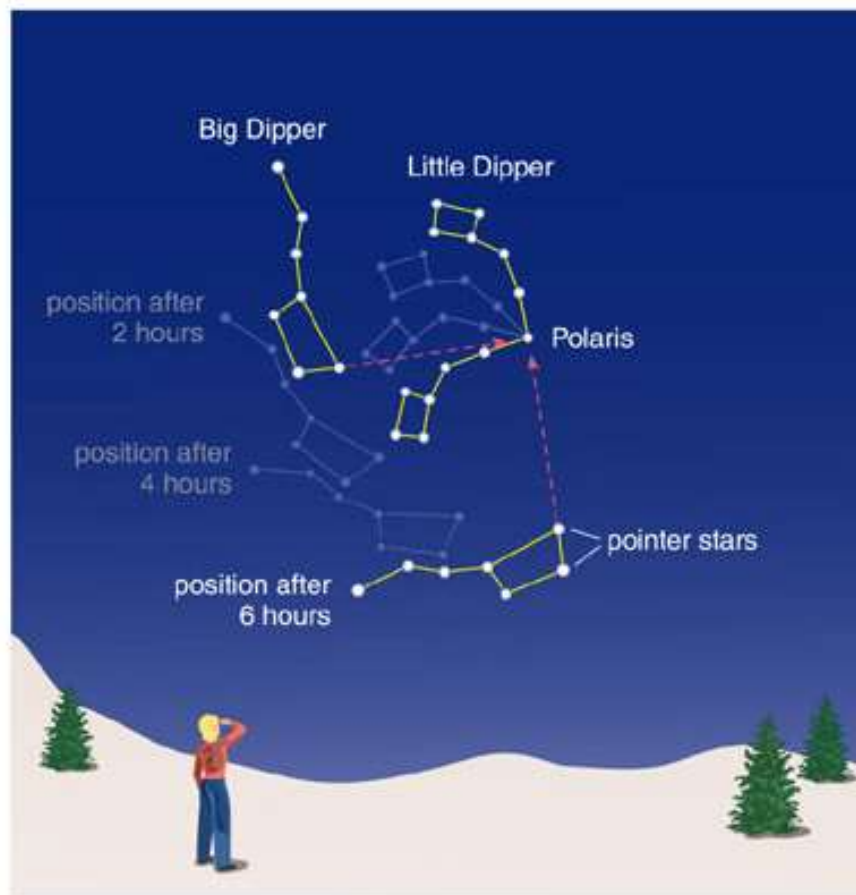


(c)

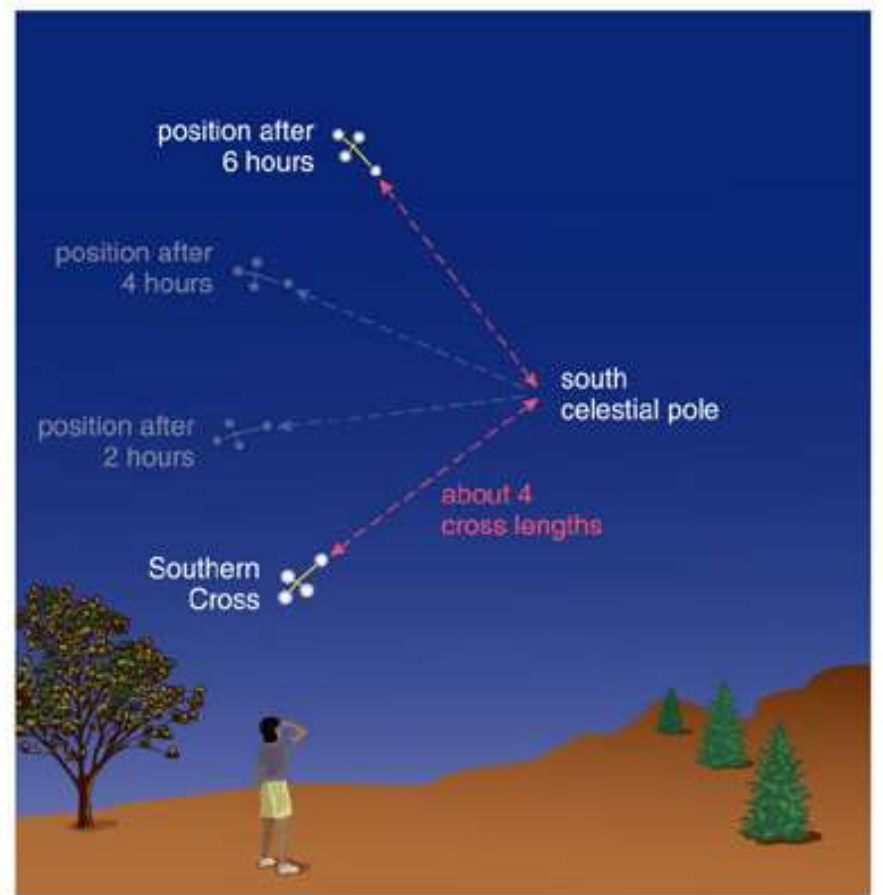
## a) At the North Pole:

- we would see only the top half of the celestial sphere.
- The Zenith would point approximately towards Polaris.
- The night sky would be almost the same each night (but for precession)
- All circumpolar stars are visible (no stars 'rise' or 'set')

# North & South Celestial poles are ~Fixed (*precession slowly changes this*) *North and South rotate in opposite directions*



looking northward



looking southward



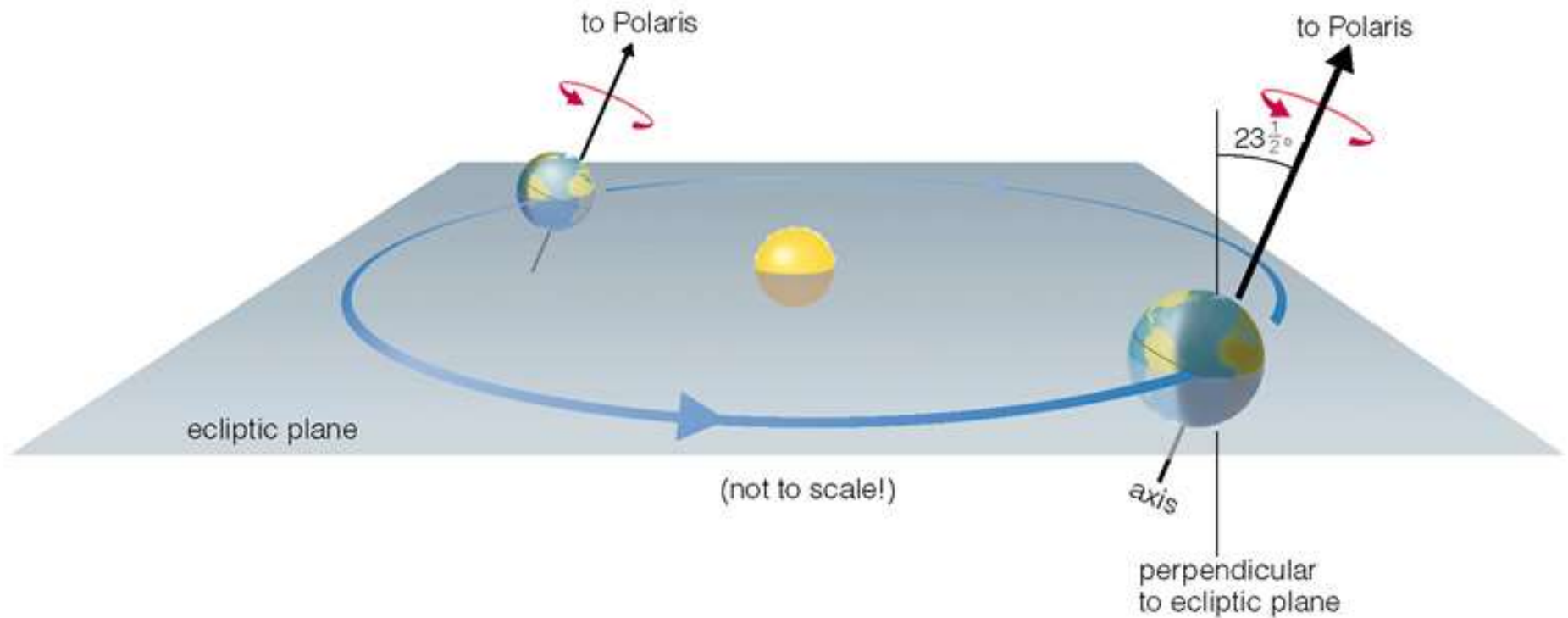


created with Advanced Stacker PLUS  
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Steven Christenson  
StarCircleAcademy.com

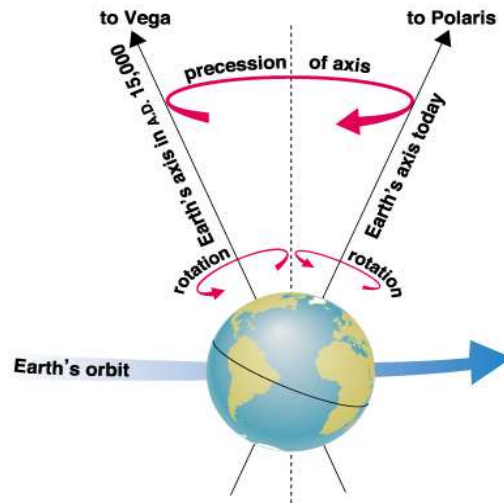
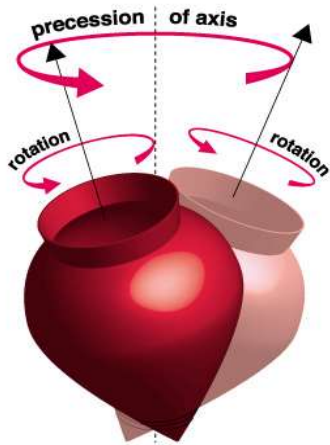


If a Star over a celestial pole is far away enough (Polaris is 434 light years away), the star position is practically constant throughout the day, and even the year



# Precession of the Equinoxes

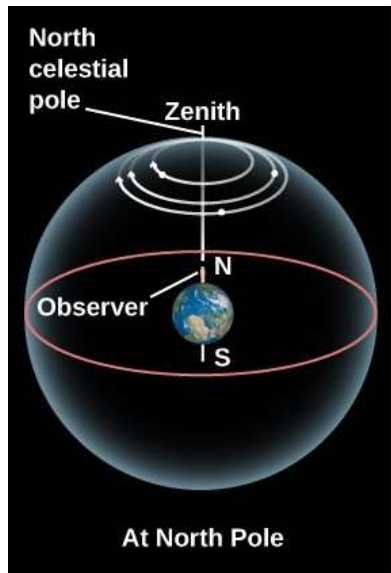
<https://www.youtube.com/watch?v=qIVgEoZDjok>



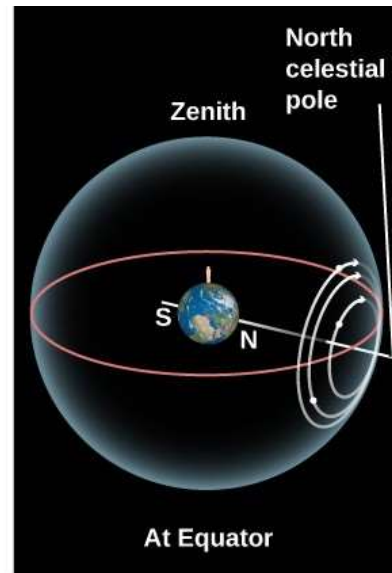
Copyright © Addison Wesley

- The Earth's axis precesses (wobbles) like a top, once about every 26,000 years.
- Precession changes the positions in the sky of the celestial poles and the equinoxes.  
⇒ *Polaris* won't always be the north star.  
⇒ The spring equinox, seen by ancient Greeks in *Aries*, moves westward and is now in *Pisces*!

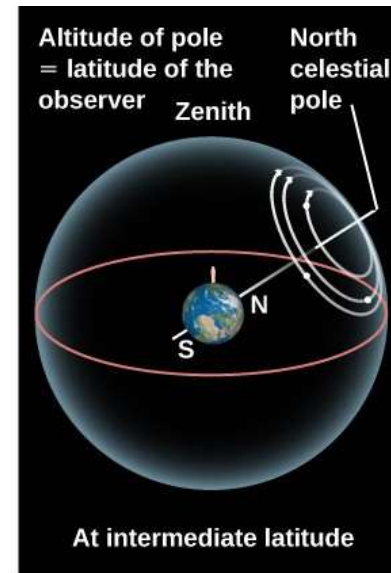
# How does the night sky change by location on the Earth?



(a)



(b)



(c)

## b) At the Equator:

- All celestial poles would lie at the horizons to the North and South
- All stars will rise and set
- There would be no circumpolar stars

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## Question #3

From what location on Earth would all of the stars on the celestial sphere be visible?

- A: The Equator
- B: The North Pole
- C: The Prime Meridian
- D: The South Pole

# iClicker2 Instructions

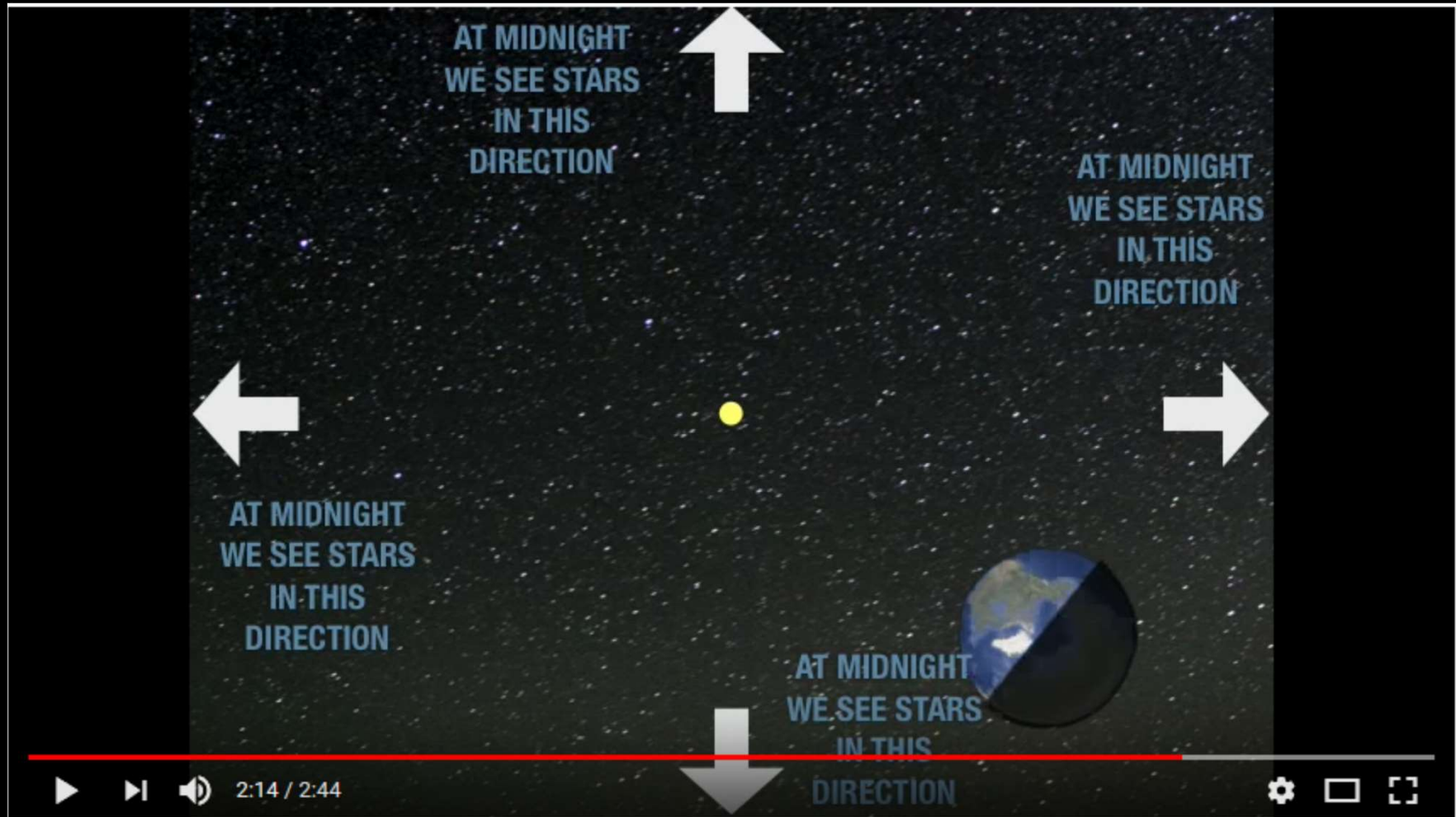
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- D: The South Pole

# At Different Times of the Year, A Different Night Sky is Visible at the Equator



<https://www.youtube.com/watch?v=kMjSiJ7V2Xc>

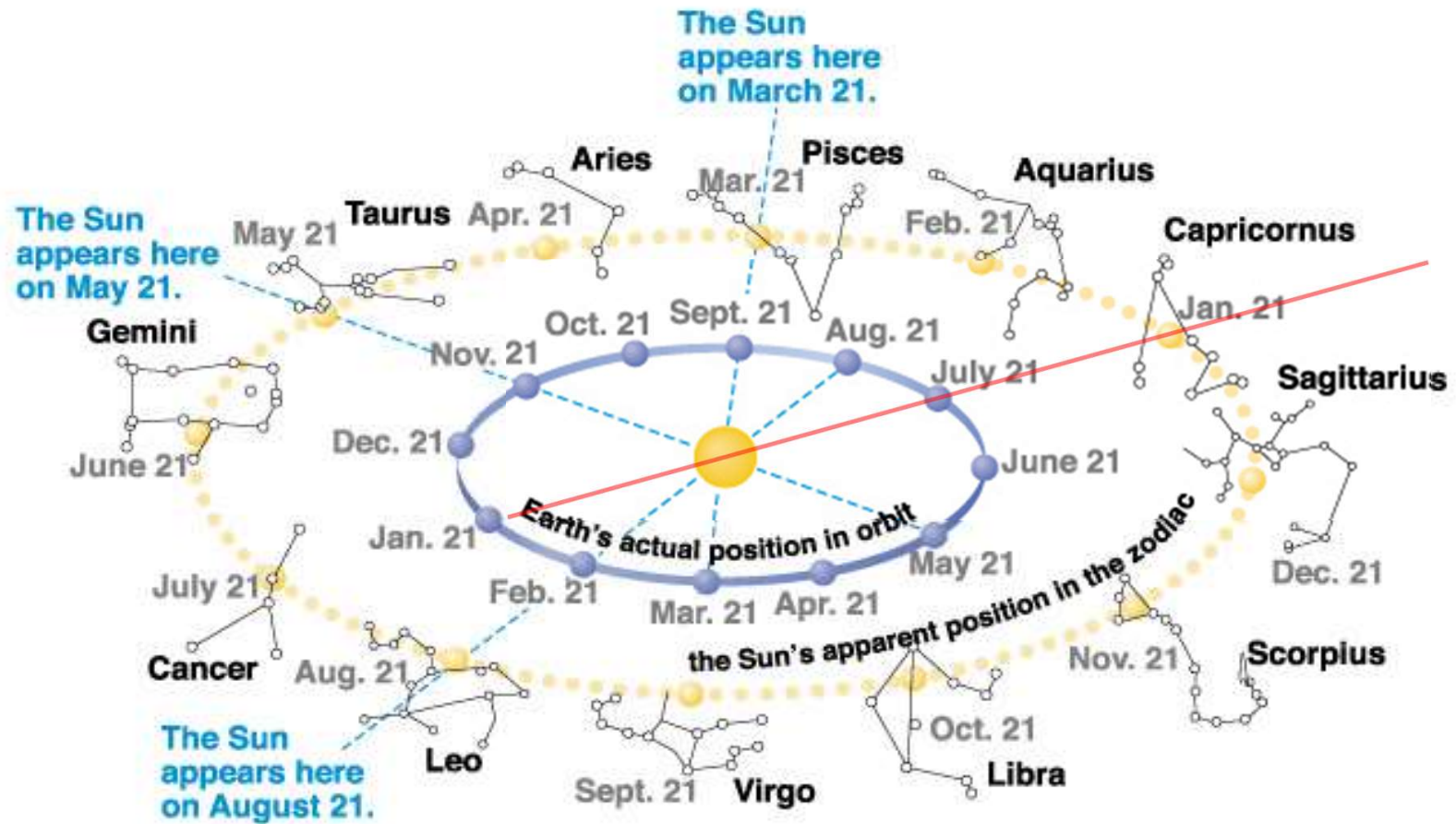
# The Constellations of the Zodiac

<http://astro.unl.edu/classaction/animations/coordsmotion/zodiac.html>

[https://www.youtube.com/watch?v=7lwJGHg\\_SQM](https://www.youtube.com/watch?v=7lwJGHg_SQM)

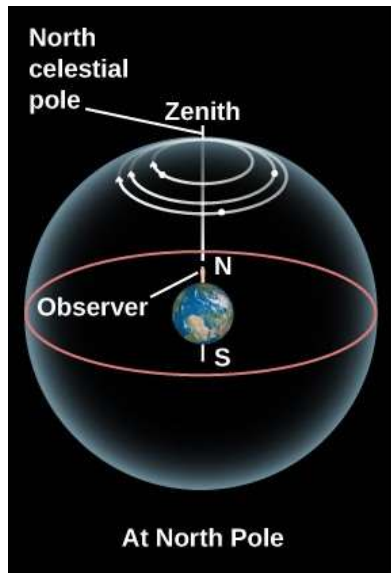


# The Sun Appears to Move Through the Constellations of the Zodiac

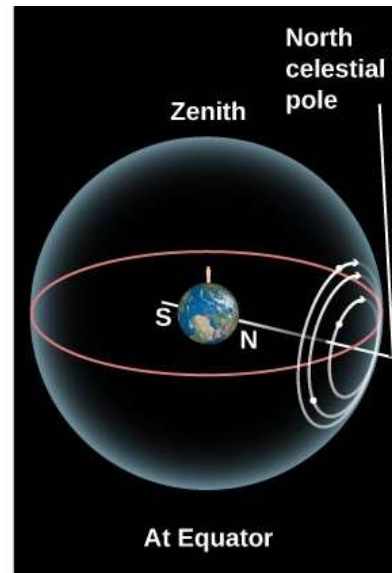




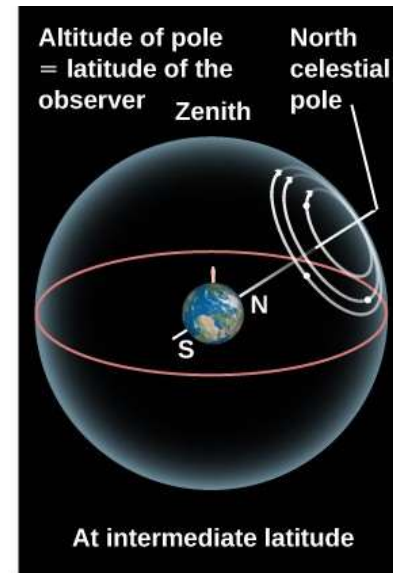
# How does the night sky change by location on the Earth?



(a)



(b)

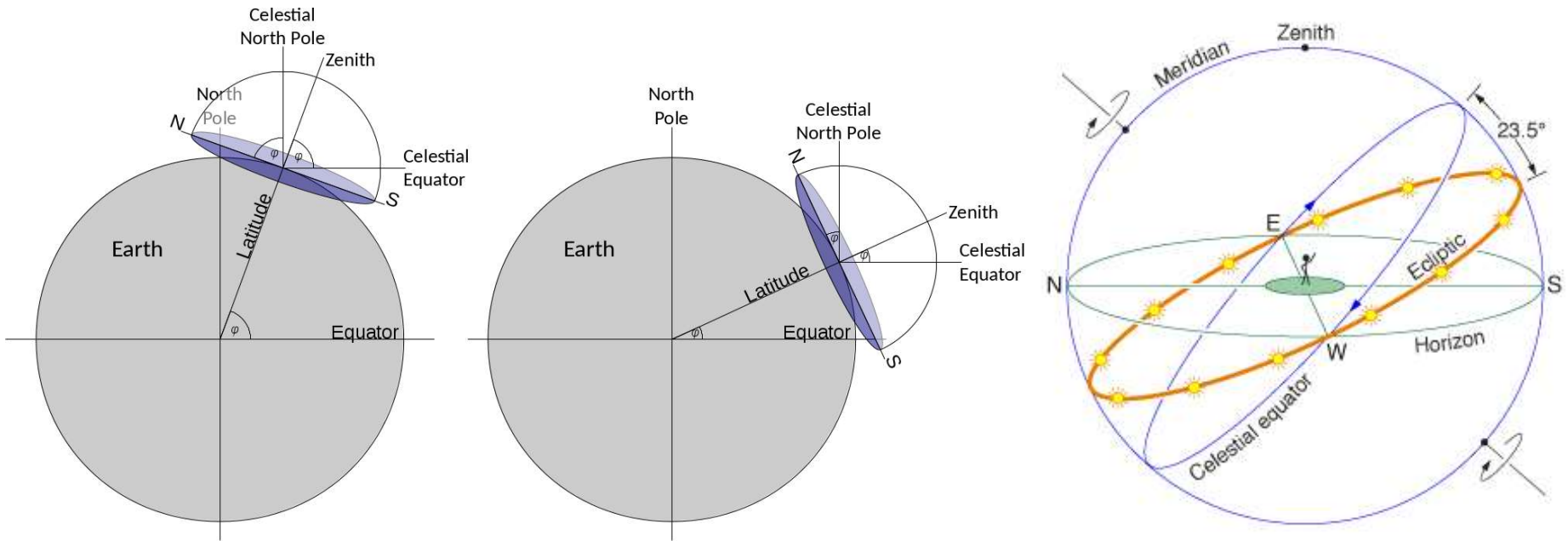


(c)

## c) At Intermediate Latitude:

- Some stars are circumpolar
- Some stars are never visible

So unless you are in one of the poles, or the equator, your local reference frame is usually *a bit* more complicated



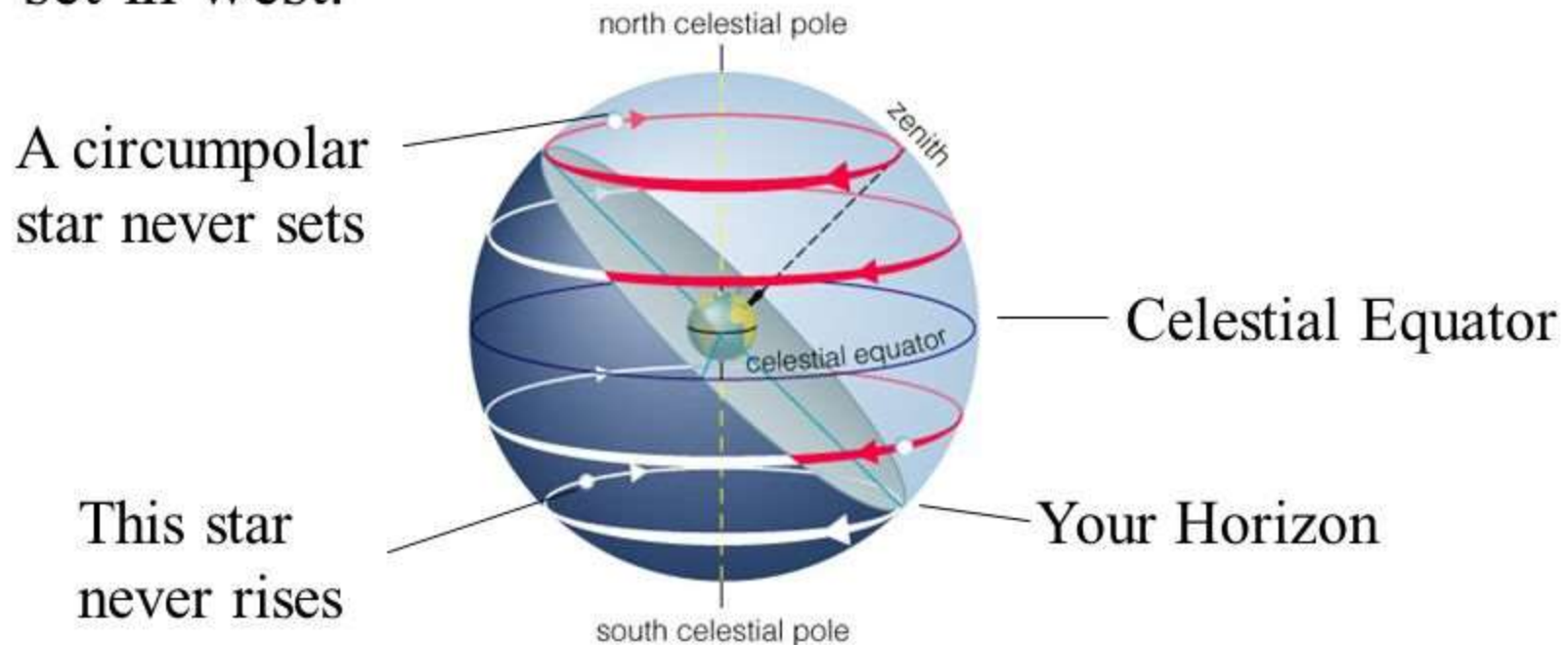
**Zenith:** lies directly above the observer

**Horizon:** plane blocked by Earth, lying  $90^\circ$  from the zenith

**Meridian:** crosses from due North over zenith to due South

# Our view from Earth:

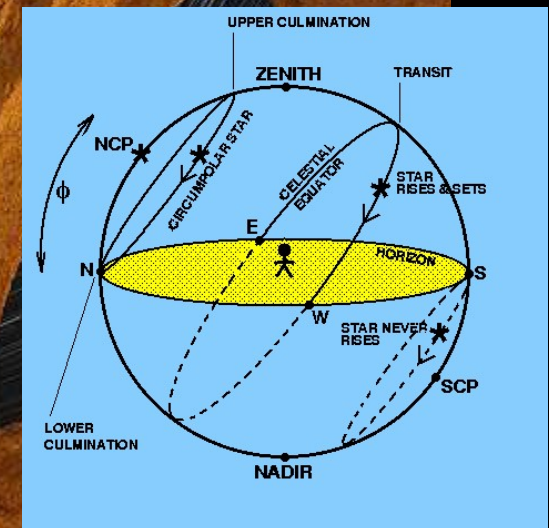
- Stars near the north celestial pole are circumpolar and never set.
- We cannot see stars near the south celestial pole.
- All other stars (and Sun, Moon, planets) rise in east and set in west.





If you are not at the North Pole, The stars will rotate around a point indicative of your latitude

- Zenith at the North pole
- ~ at the Horizon if at the Equator





<http://weside.deviantart.com>

The Seasons are determined based on how close the Earth is to the Sun

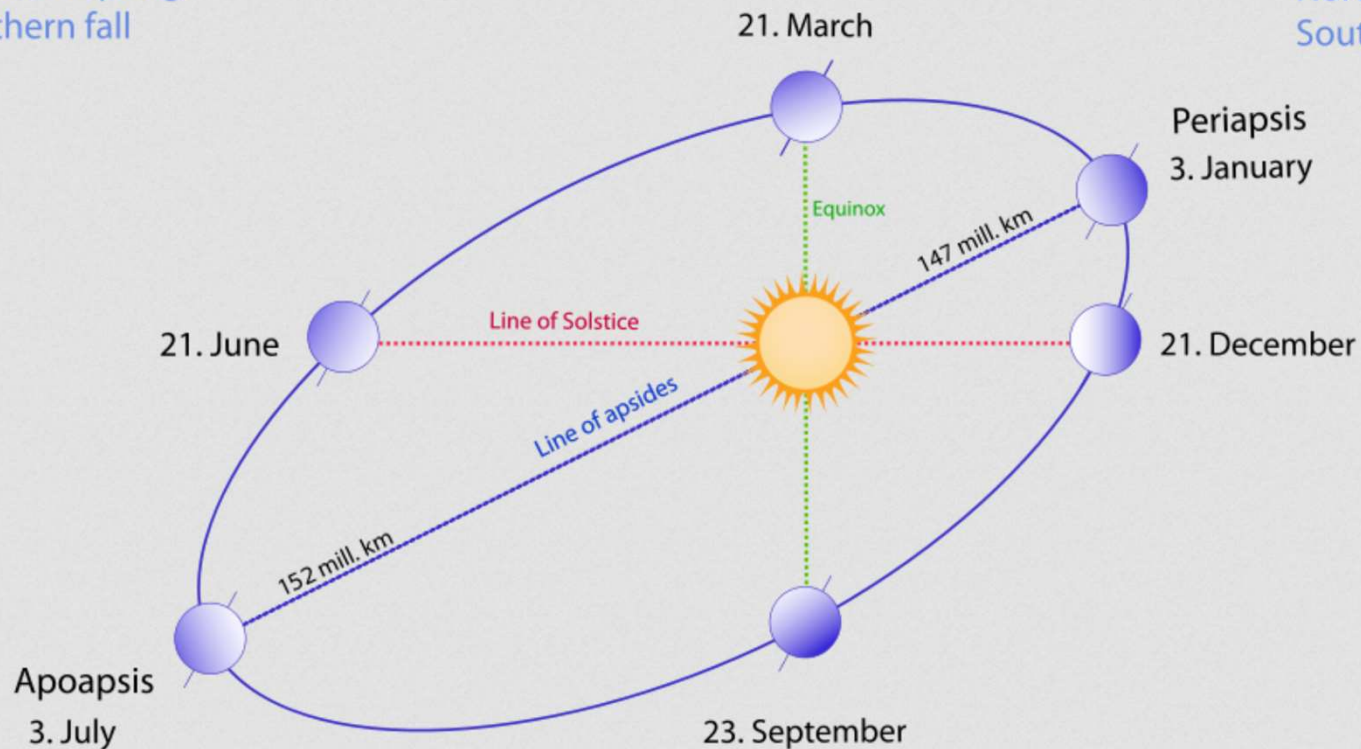
♦ A: True

♦ B: False

# Actually, for Winter in the Northern Hemisphere, the Sun is further away!

Northern spring/  
Southern fall

Northern winter/  
Southern summer

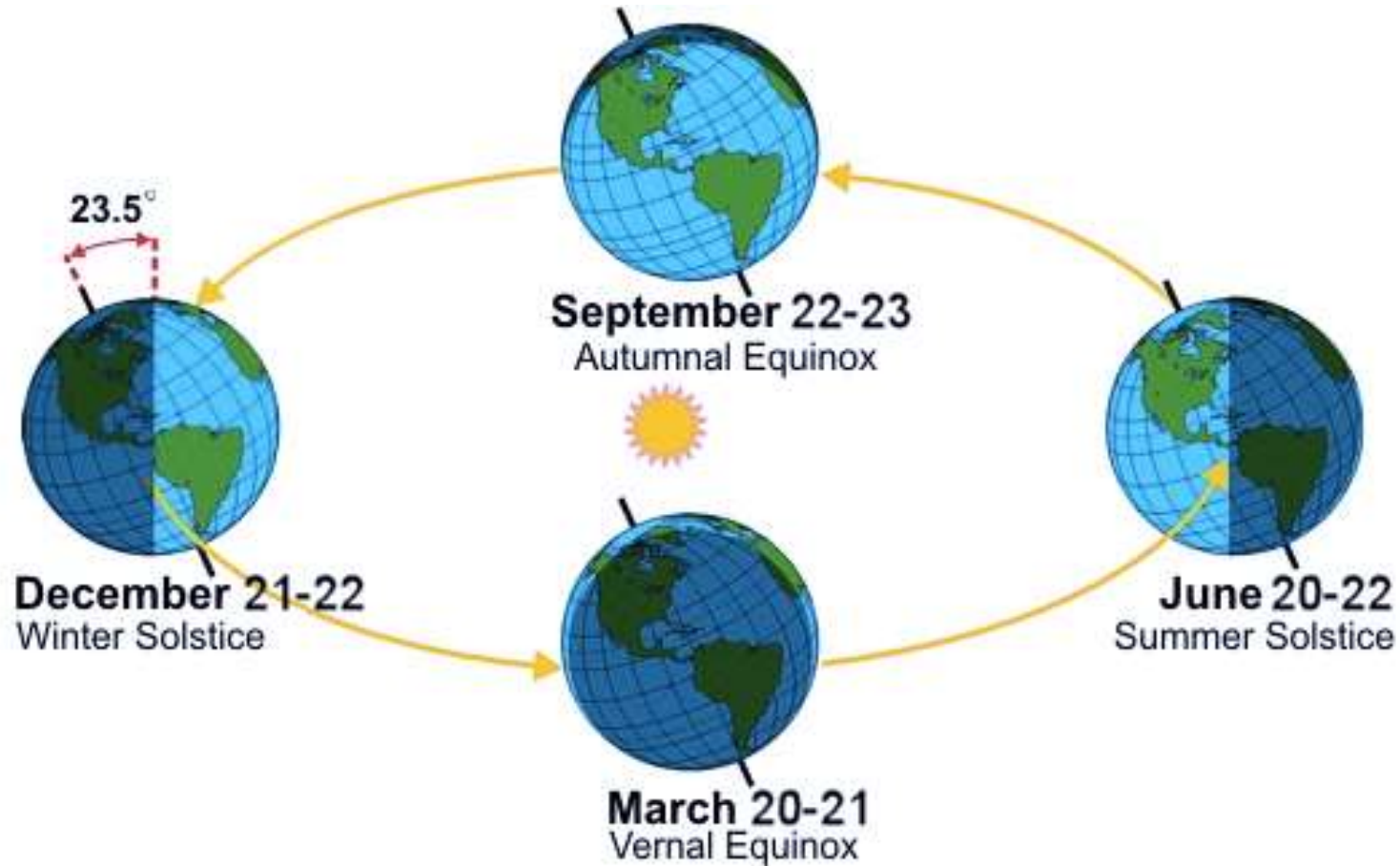


Northern summer/  
Southern winter

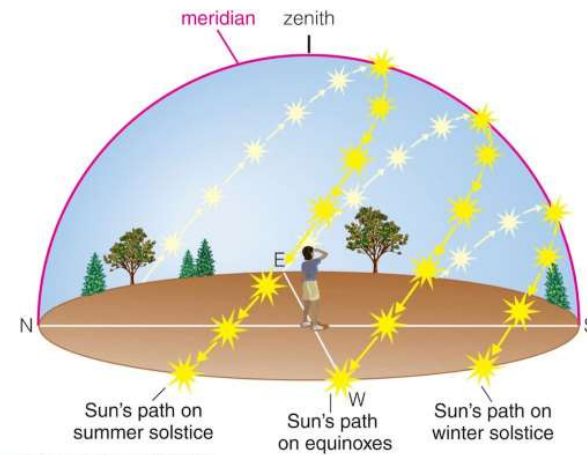
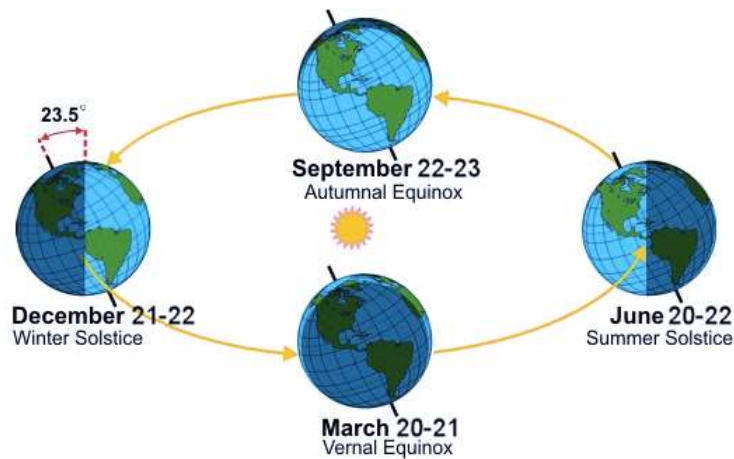
Northern fall/  
Southern spring



# The Seasons are caused by the Tilt of Earth's axis



# Extremes of Solstices & Equinoxes



**Path of the Sun on the summer solstice at the Arctic Circle**

Approximate time:  
Direction:

Midnight  
due north

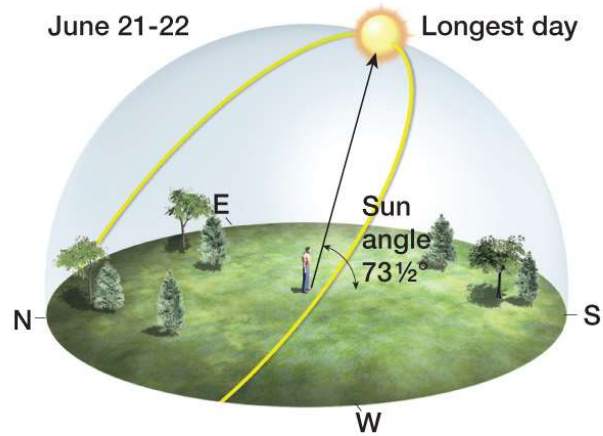
6:00 A.M.  
due east

Noon  
due south

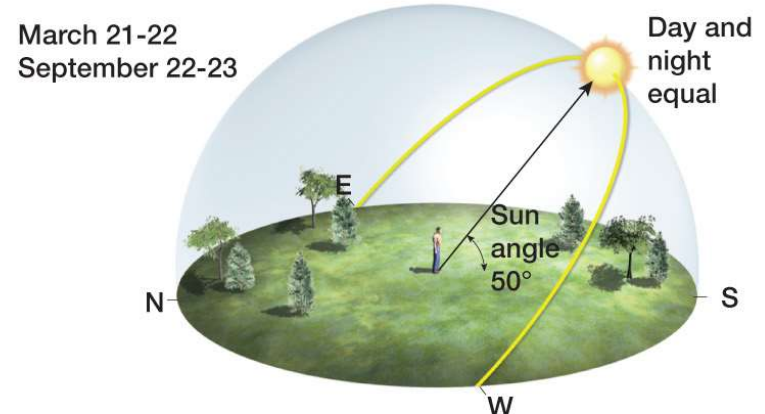
6:00 P.M.  
due west



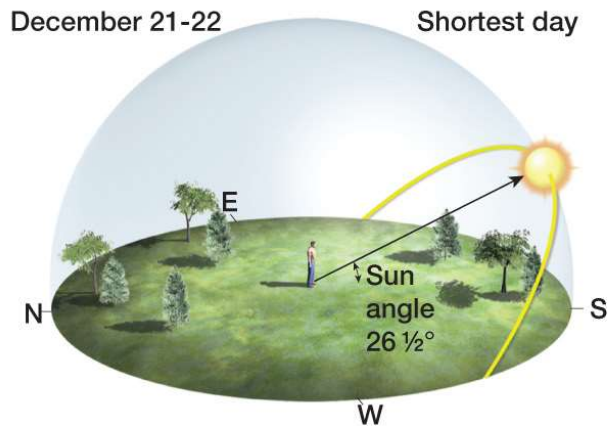
# Sun angle changes with Season



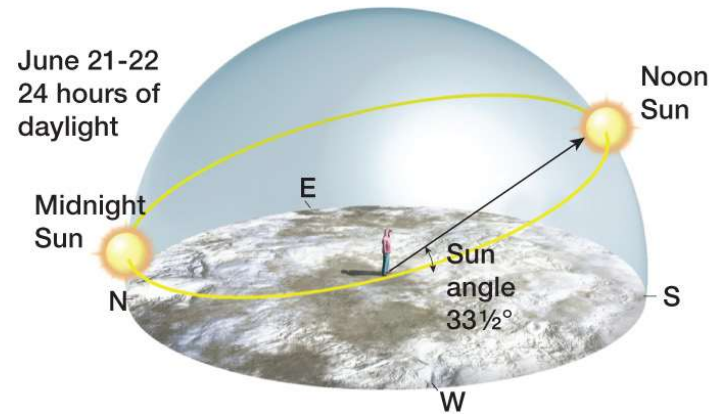
(a) Summer solstice at 40° latitude



(b) Spring or fall equinox at 40° latitude

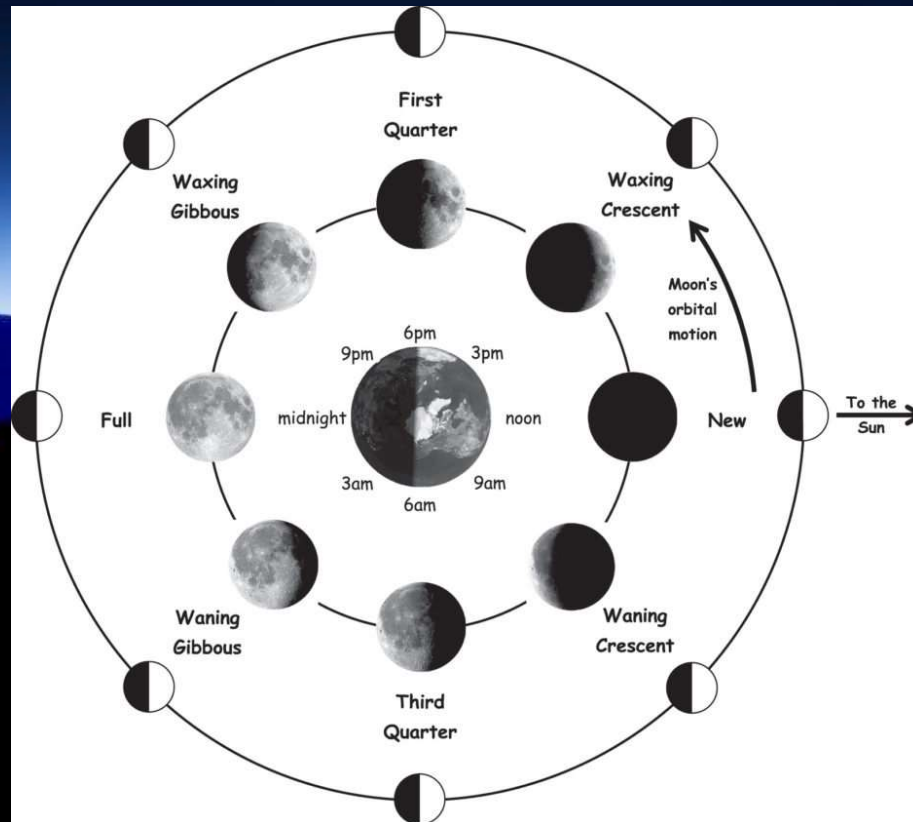


(c) Winter solstice at 40° latitude



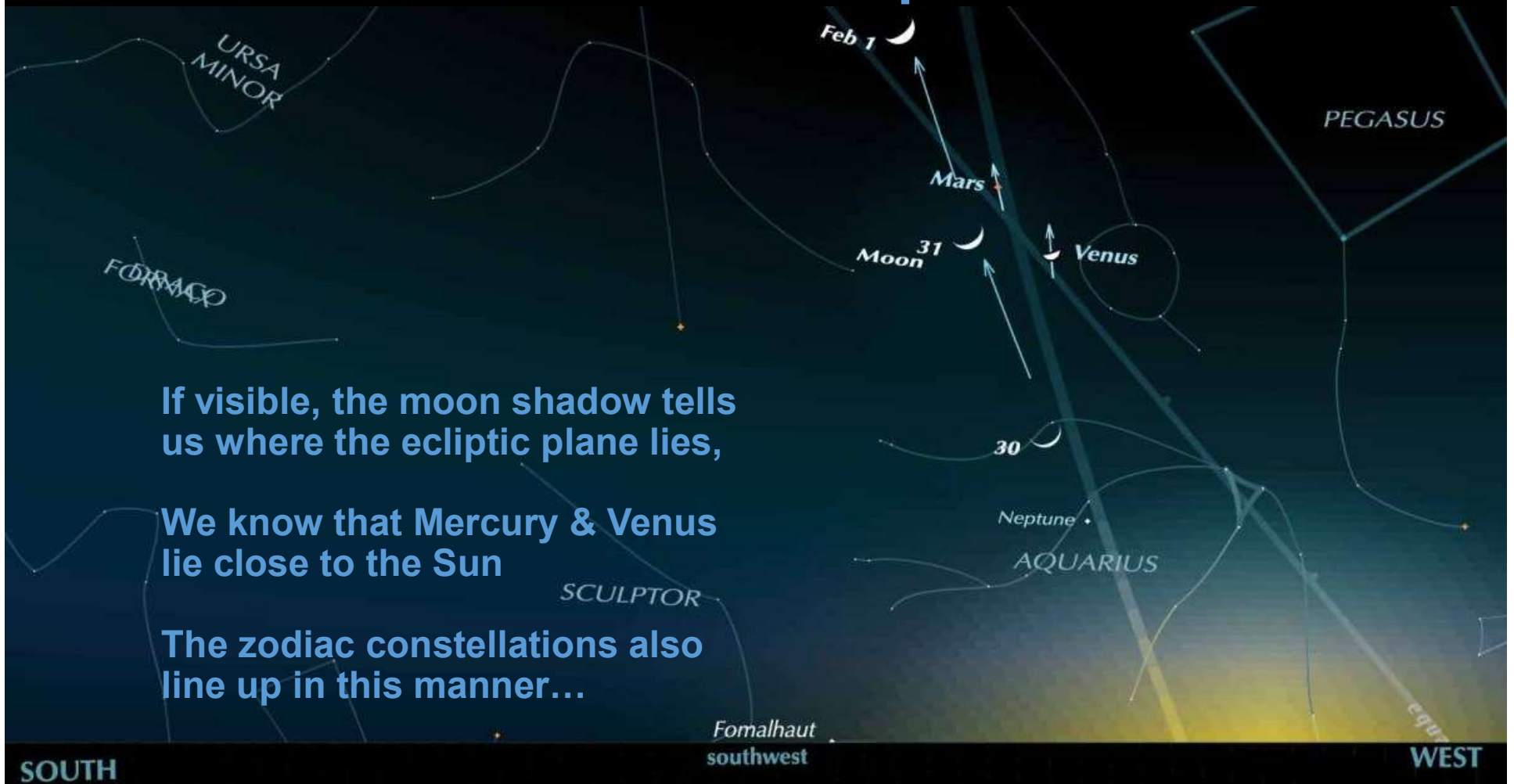
(d) Summer solstice at 80° latitude

# The Phases of the Moon



The moon orbits the Earth every ~27 days

# The Planets, Sun, Moon, Constellations of the Zodiac All Lie in the Ecliptic Plane



If visible, the moon shadow tells us where the ecliptic plane lies,

We know that Mercury & Venus lie close to the Sun

The zodiac constellations also line up in this manner...

2017 Jan 31 Tue, 1 hour after sunset  
6:19 PM CST = Feb 1, 0:19 UT  
latitude 40°N, longitude 90°W  
sidereal time 46° = 3.08h

↑ Sun

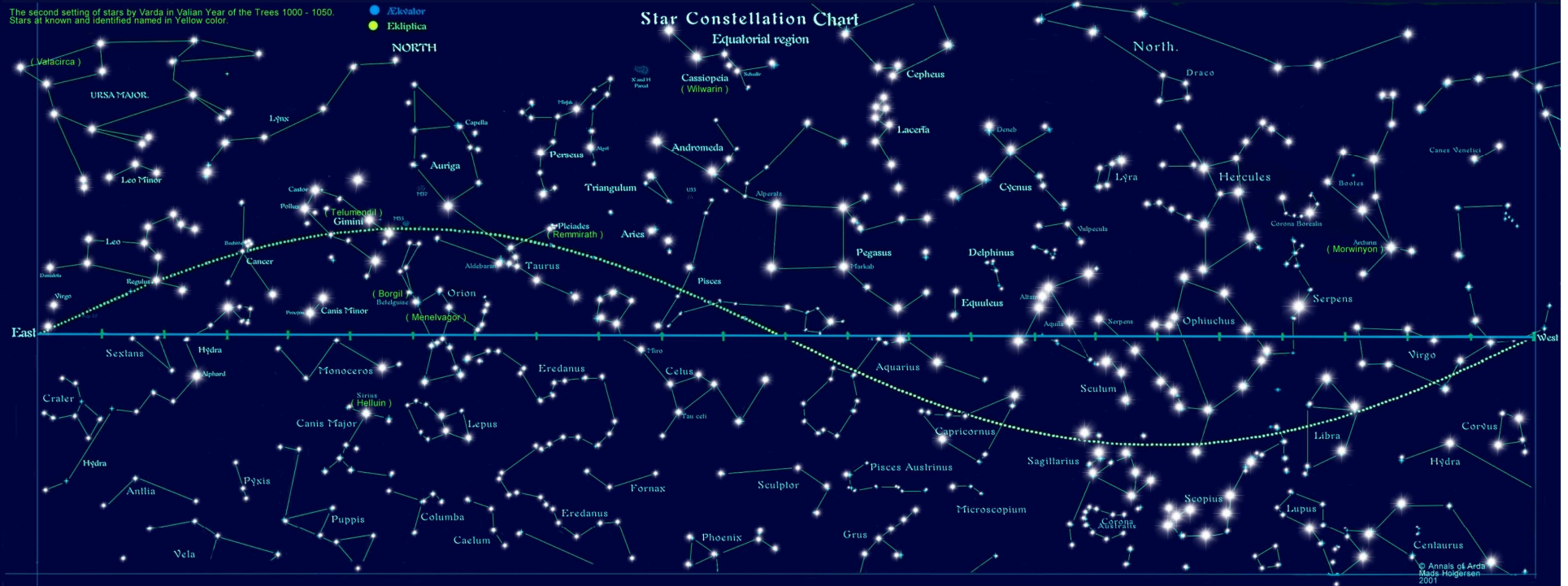
# The 88 Constellations

Begin at East (Vernal or Spring Equinox)

Here, Equator is shown in Blue horizontal line

Ecliptic plane as it changes with seasons is green

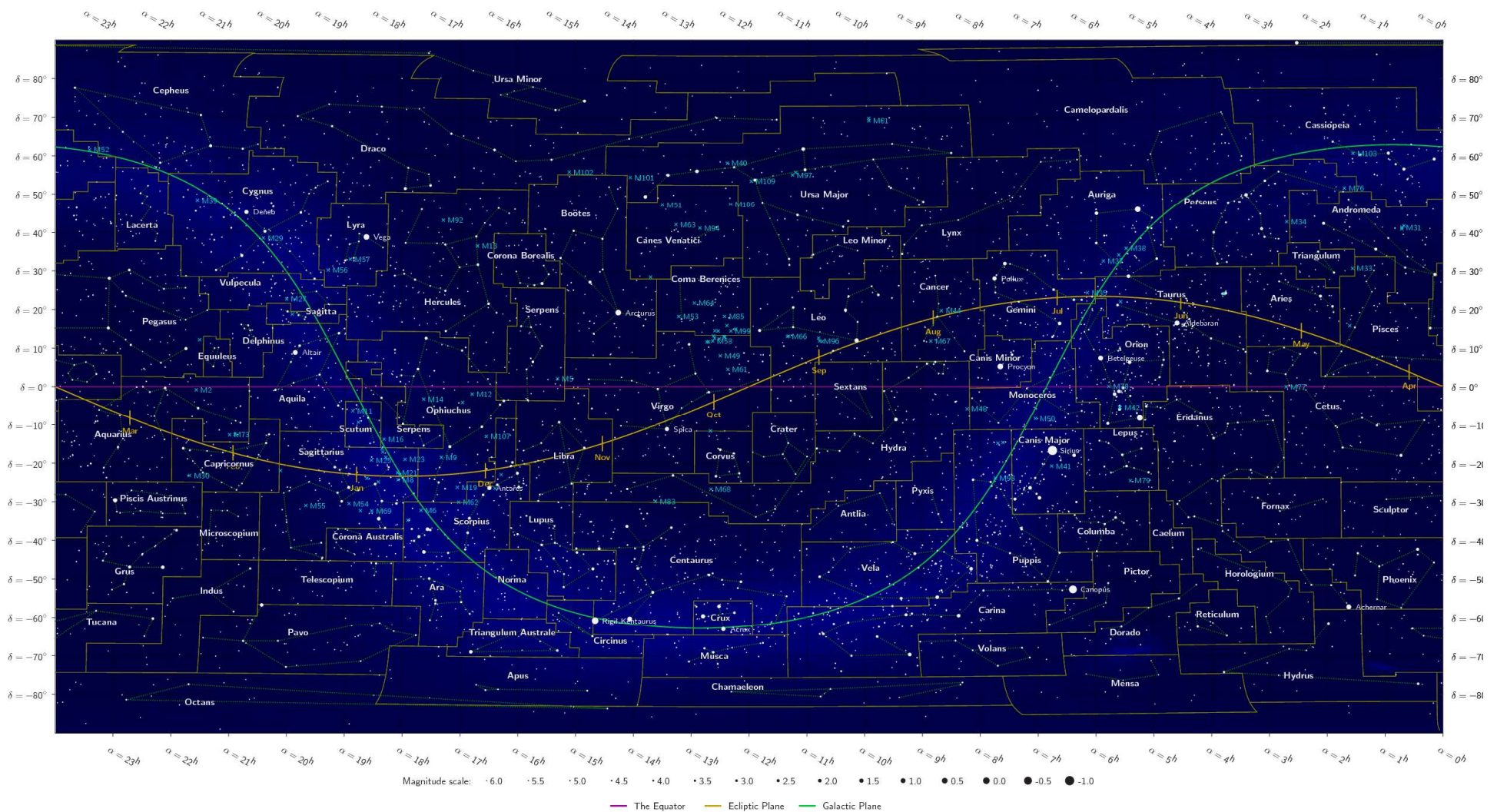
## 12 HOROSCOPES



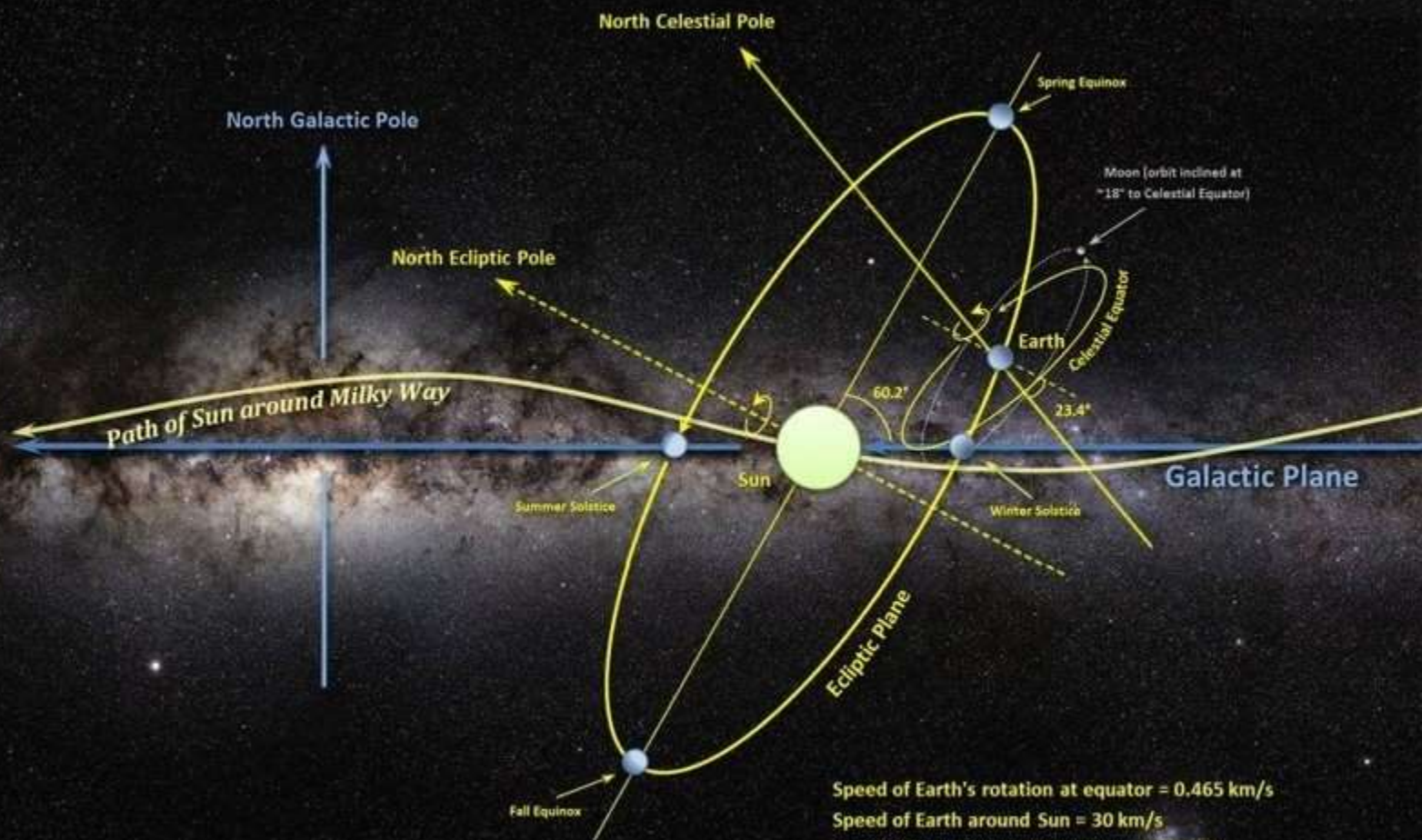


# What About the Galactic Plane?

[https://in-the-sky.org/data/constellations\\_map\\_equ11012.png](https://in-the-sky.org/data/constellations_map_equ11012.png)



## MOTION OF EARTH AND SUN AROUND THE MILKY WAY



Speed of Earth's rotation at equator = 0,465 km/s  
Speed of Earth around Sun = 30 km/s  
Speed of Sun around Milky Way = 230 km/s  
Sun is approximately 26,000 light years from Galactic Center

*Diagram Not to Scale*

End of Today's Lecture